



## Slow subduction of a thick ultrahigh-pressure terrane

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[1] High-precision Lu-Hf and Sm-Nd ages reported here cover a large portion of the Western Gneiss Region (WGR) ultrahigh-pressure (UHP) terrane, Norway, and range from  $413.9 \pm 3.7$  to  $397.1 \pm 4.8$  Ma. Collectively, new and existing geochronologic data for eclogites demonstrate that eclogite facies metamorphism occurred over a large area ( $60,000 \text{ km}^2$ ) for an unexpectedly long period of  $>20$  Ma. A thermal model using slow subduction ( $2\text{--}4 \text{ mm a}^{-1}$ ) of a relatively warm and thick slab can reproduce the observed  $P$ - $T$ - $t$  history of the WGR UHP terrane, indicating that large UHP terranes with protracted peak or near-peak histories probably reflect slow subduction of relatively thick crustal sections. **Citation:** Kylander-Clark, A. R. C., B. R. Hacker, C. M. Johnson, B. L. Beard, and N. J. Mahlen (2009), Slow subduction of a thick ultrahigh-pressure terrane, *Tectonics*, 28, TC2003, doi:10.1029/2007TC002251.

### 1. Introduction

[2] The number of recognized ultrahigh-pressure (UHP) rocks worldwide has increased dramatically in the last two decades, and as such, awareness of their importance in processes such as orogenesis, the generation and rearrangement of continental crust, and interaction between the crust and the mantle has grown. Significant advances have been made toward understanding the evolution of UHP terranes, but some fundamental processes remain enigmatic, particularly their rates and mechanisms of formation.

[3] UHP terranes are distinguished by the presence of coesite, and the best studied UHP terranes appear to have formed by subduction to depths of  $100\text{--}135 \text{ km}$  [Chopin *et al.*, 1991; Schertl *et al.*, 1991; Carswell *et al.*, 1997; Reinecke, 1998; Cuthbert *et al.*, 2000; Masago, 2000; Ota *et al.*, 2000; Carswell and Cuthbert, 2003; Hirajima and Nakamura, 2003; Krogh Ravn and Terry, 2004]. With rare exceptions, UHP rocks show peak temperatures of  $600\text{--}800^\circ\text{C}$  [Hacker, 2006]. These observations have led to the paradigm that UHP terranes were not heated to higher

temperatures because they were rapidly subducted, experience peak  $P$ - $T$  conditions for a short time [e.g., Carswell *et al.*, 2003b; Zheng *et al.*, 2003], and were then rapidly exhumed and cooled [e.g., Eide *et al.*, 1994; Rubatto and Hermann, 2001; Hacker *et al.*, 2003; Root *et al.*, 2005; Parrish *et al.*, 2006]. Geochronologic studies suggest that some UHP terranes, indeed, spent only a few million years at HP conditions [e.g., Leech *et al.*, 2005]. However, other UHP terranes in eastern China [Hacker *et al.*, 2006; Liu *et al.*, 2006a], western China [Mattinson *et al.*, 2006], and Greenland [McClelland *et al.*, 2006] were held at high-pressure depths for  $15\text{--}20$  Ma. The first two of these are so-called giant UHP terranes (more than a few thousand square kilometers), which might well have evolved over a more protracted period than smaller UHP terranes.

[4] Geochronologic studies also show that some UHP terranes underwent rapid near-isothermal decompression, followed by rapid cooling [Maruyama and Parkinson, 2000; Liu *et al.*, 2006b; Parrish *et al.*, 2006; Hacker, 2007]. The need for rapid cooling has led to a second paradigm, that UHP terranes are thin ( $<10 \text{ km}$  [e.g., Ernst, 2006]), although field relations and thermal models have been used to argue that some UHP terranes are thick [Hacker *et al.*, 2000; Root *et al.*, 2005].

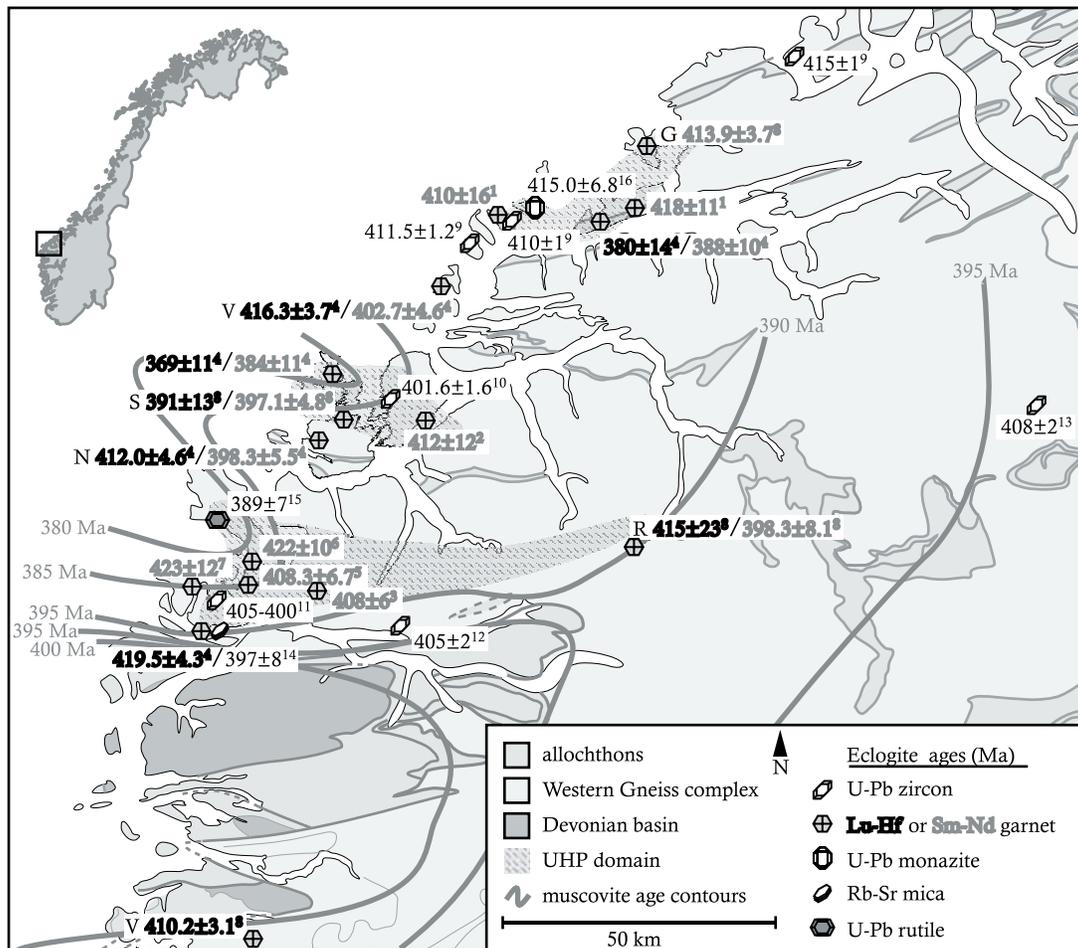
[5] In this paper we present new Sm-Nd and Lu-Hf ages that bolster earlier indications [Kylander-Clark *et al.*, 2007] of a 20-Ma duration for (U)HP metamorphism in western Norway. We then use simple thermal models to assess the tectonic settings under which protracted prograde metamorphism, followed by rapid exhumation and cooling, is possible. These results bear on questions such as the following: What dimensions can a UHP body have if it is to cool fast enough to retain HP mineral assemblages and yield the temperature-time curves observed in UHP terranes? What subduction velocities are possible? What changes in subduction behavior are permissible?

### 2. Geologic Background

[6] The WGR consists chiefly of tonalitic to granodioritic Baltica gneiss, called the Western Gneiss Complex (WGC), which is overlain by a stack of oceanic and continental allochthons (Figure 1). The allochthons are principally  $1700\text{--}950 \text{ Ma}$  [Corfu, 1980; Schärer, 1980; Tucker *et al.*, 1990; Root *et al.*, 2005] and were emplaced eastward over the WGC during the Scandian Orogeny at  $\sim 435\text{--}400 \text{ Ma}$  [Roberts and Gee, 1985; Hacker and Gans, 2005]. It was during this orogeny that the signature UHP metamorphism of the Western Gneiss Region developed at  $P$ - $T$  conditions

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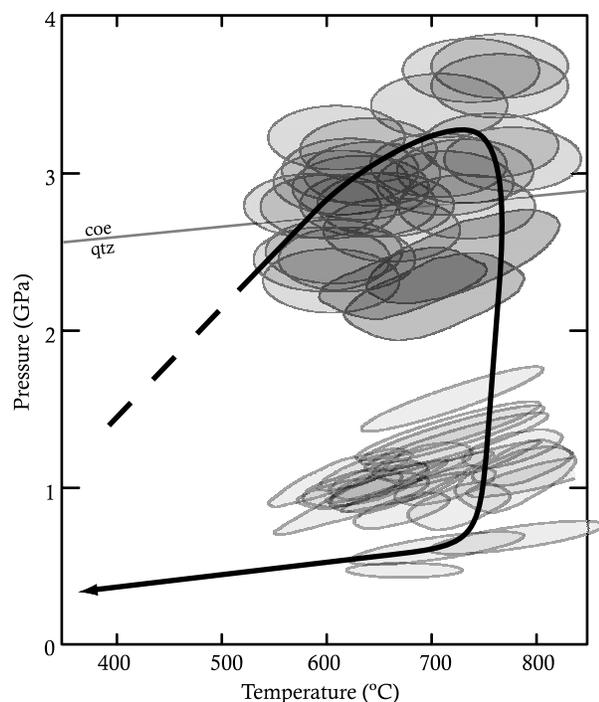


**Figure 1.** Western Gneiss Region, western Norway, showing the following eclogite ages, Sm-Nd and Lu-Hf: Values with superscript numbers are from 1, *Mørk and Mearns* [1986]; 2, *Jamtveit et al.* [1991]; 3, *Mearns* [1986]; 4, *Kylander-Clark et al.* [2007] (V, Vigra; N, NW Gurskøy); 5, *Carswell et al.* [2003a]; 6, H. K. Brueckner (personal communication, 2007); 7, *Griffin and Brueckner* [1980]; 8, this study (G, Gossa; R, Geiranger; V, Vollstein). U-Pb zircon, 9, *Krogh et al.* [2004]; 10, *Carswell et al.* [2003b]; 11, *Root et al.* [2004]; 12, *Young et al.* [2007b]; 13, *Kylander-Clark et al.* [2007]. Rb-Sr, 14, *Griffin and Brueckner* [1985]. U-Pb rut-omph, 15, *Schärer and Labrousse* [2003]. U-Pb monazite, 16, *Terry et al.* [2000a]. The  $^{40}\text{Ar}/^{39}\text{Ar}$  K-white mica age contours, *Hacker et al.* [2006].

as high as 800°C and 3.4 GPa (Figure 2) [e.g., *Griffin et al.*, 1985; *Andersen et al.*, 1991; *Cuthbert et al.*, 2000]. UHP eclogites are found along the Atlantic coast in three distinct domains (Figure 1) [*Root et al.*, 2005], the northernmost of which contains the highest peak pressures and temperatures [*Dobrzhinetskaya et al.*, 1995; *van Roermund and Drury*, 1998; *Terry et al.*, 2000b; *van Roermund et al.*, 2001; *Carswell et al.*, 2006]. Metamorphic grade and peak pressures decrease eastward across the Western Gneiss Region toward the foreland, such that the easternmost reported eclogite is more than 150 km east of the coast. This metamorphic gradient is interpreted to reflect down-to-the-west subduction. The peak UHP metamorphism was followed by a pervasive amphibolite facies overprint from pressures of ~1.5 GPa down to 0.5 GPa (~45 to 15 km)

during near-isothermal exhumation to midcrustal depths (summary by *Hacker* [2006]).

[7] Attempts to determine the age of the UHP metamorphism began with *Griffin and Brueckner's* [1980] pioneering Sm-Nd work on eclogites (Figure 3). They reported a broad range of ages, 447–407 Ma, from which they selected an average age of 425 Ma. Subsequent Sm-Nd work (*Mearns* [1986] (recalculated by R. B. Root), *Mørk and Mearns* [1986], and *Carswell et al.* [2003a]) produced only younger ages, 412–408 Ma. More recent work on the U-Pb system in zircon gave ages of 415–400 Ma [e.g., *Carswell et al.*, 2003b; *Krogh et al.*, 2004; *Root et al.*, 2004; *Young et al.*, 2007a], and *Root et al.* [2004] drew on these data to conclude that the older Sm-Nd ages were mixed ages that were contaminated by older cores that predate the UHP event. Our Sm-Nd ages, as well as the first Lu-Hf ages from



**Figure 2.** *P-T* history of the Western Gneiss Region after Hacker [2006], showing peak conditions of  $\sim 750^{\circ}\text{C}$  and 3.4 GPa, followed by near-isothermal decompression to midcrustal levels; coe, coesite; qtz, quartz.

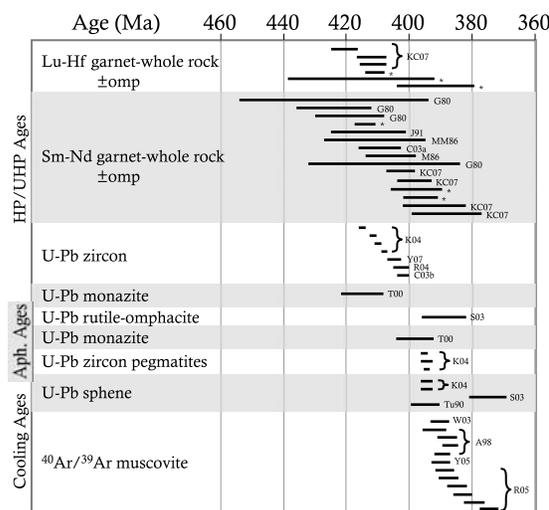
Norway, span a large range of  $419.5 \pm 4.3$  to  $369 \pm 11$  Ma [Kylander-Clark *et al.*, 2007], and these have been interpreted to confirm a broad age range for eclogite facies metamorphism. The age of the amphibolite facies overprint is constrained by  $\sim 395$ – $390$  Ma titanite, zircon and monazite ages [Tucker *et al.*, 1990; Terry *et al.*, 2000a; Krogh *et al.*, 2004; Tucker *et al.*, 2004; Kylander-Clark *et al.*, 2008], and  $^{40}\text{Ar}/^{39}\text{Ar}$  muscovite ages of 400 Ma to 374 Ma (summary by Hacker [2006]). These data require exhumation from UHP depths of 120 km to midcrustal depth of 15 km in  $\sim 5$ – $15$  Ma.

### 3. New Lu-Hf and Sm-Nd Ages

[8] The Lu-Hf and Sm-Nd ages we previously reported [Kylander-Clark *et al.*, 2007] were obtained from a relatively restricted portion of the WGR within and around the UHP domains. Here we report ages from four eclogite localities that were selected to broaden the areal coverage to test the proposal that protracted UHP metamorphism occurred over a broad region. The combination of Lu-Hf and Sm-Nd ages from the same sample separates is advantageous, as it yields a minimum duration of garnet stability for any given sample, and thus provides a direct constraint on the duration of metamorphism. Because the distribution coefficient for Lu in garnet is relatively high with respect to other major eclogite phases, such as omphacite [McKenzie and O’Nions, 1991], and there are no minor phases that control the budget of Lu, the bulk of the Lu is sequestered in

the core during garnet growth. Thus, unless the temperature exceeded the blocking temperature for the Lu-Hf system, Lu-Hf ages are biased toward early garnet growth, and this generalization holds true over a wide range of garnet growth mechanisms [Skora *et al.*, 2006]. The distribution coefficient of Sm in garnet, on the other hand, is much less than that of Lu. Thus, unless the temperature exceeded the blocking temperature for the Sm-Nd system, Sm-Nd ages tend to represent the average or later stages of garnet growth. Therefore, the difference between the Lu-Hf and Sm-Nd ages for samples can be used as a minimum estimate for the duration of garnet growth, i.e., the duration of HP and UHP conditions, provided that the samples never significantly exceeded the blocking temperature for either system. This appears to be the case in the WGR: The Sm-Nd garnet ages of eclogites are equivalent to U-Pb zircon ages that are interpreted to represent the eclogite facies metamorphism of nearby sample localities (see section 3.3).

[9] Sample preparation and analysis followed that outlined by Kylander-Clark *et al.* [2007] and Lapen *et al.* [2003, 2004]. Analyses were performed on a Micromass IsoProbe ICP-MS (Lu-Hf) and VG Sector 54 TIMS (Sm-Nd) at the University of Wisconsin, Madison. Pressure and temperature estimates were obtained from minerals analyzed on the Cameca SX-50 microprobe at the University of



**Figure 3.** Geochronology from the Western Gneiss Region, showing ages of prograde through peak conditions, postpeak amphibolite facies metamorphism, and cooling; omp, omphacite; Aph, amphibolite. KC07, Kylander-Clark *et al.* [2007]; MM86, Mørk and Mearns [1986]; J91, Jamtveit *et al.* [1991]; M86, Mearns [1986]; C03a, Carswell *et al.* [2003a]; G80, Griffin and Brueckner [1980]; K04, Krogh *et al.* [2004]; C03b, Carswell *et al.* [2003b]; R04, Root *et al.* [2004]; Y07, Young *et al.* [2007b]; S03, Schärer and Labrousse [2003]; Tu90, Tucker *et al.* [1990]; T00, Terry *et al.* [2000a]; W03, Walsh [2003]; A98, Andersen *et al.* [1998]; Y05, Young [2005]; R05, Root *et al.* [2005]; asterisk, this study.

**Table 1a.** Lutetium-Hafnium Isotope Data for Eclogites From the Western Gneiss Region, Norway<sup>a</sup>

Sample <sup>b</sup>	Concentration (ppm)		<sup>176</sup> Lu/ <sup>177</sup> Hf	<sup>176</sup> Hf/ <sup>177</sup> Hf	Age (Ma)	$\epsilon_{\text{Hf}(0)}$ <sup>c</sup>
	Lu	Hf				
K5622A2 (Vollstein)						
grt (107.8)	0.23	0.10	0.3255	0.285356 ± 13		
cpx (29.6)	0.00	0.24	0.0010	0.282868 ± 19		
wr (91.8)	0.14	0.42	0.0485	0.283223 ± 10	410.2 ± 3.1	12
P5701A (Sandvik)						
grt <sup>d</sup> (95.4)	0.42	0.83	0.0719	0.282914 ± 4		
grt <sup>e</sup> (86.8)	0.42	1.05	0.0568	0.282819 ± 5		
cpx (53.3)	0.01	1.01	0.0009	0.282396 ± 5		
wr (44.7)	0.20	1.66	0.0173	0.282527 ± 4	391 ± 13	-4.7
E1612Q5 (Geiranger)						
grt <sup>3</sup> (70.7)	0.53	1.23	0.0617	0.282828 ± 5		
grt <sup>d</sup> (93.4)	0.54	1.36	0.0565	0.282817 ± 3		
cpx (37.0)	0.00	0.52	0.0013	0.282369 ± 7		
wr (183.4)	0.30	2.25	0.0192	0.282518 ± 3	415 ± 23	-5.3

<sup>a</sup>Analysis accomplished by isotope dilution multicollector plasma mass spectrometry at University of Wisconsin, Madison. Decay constant used for <sup>176</sup>Lu is  $1.865 \times 10^{-11} \text{ a}^{-1}$  [Scherer *et al.*, 2001]. Errors in Hf isotope ratios are expressed as 2SE from internal measurements and refer to least significant digits. Isochron ages were calculated using Isoplot v. 3.0. Errors calculated for ages are based solely on external reproducibility of spiked standards and whole rock samples (errors for individual analyses are negligible); <sup>176</sup>Lu/<sup>177</sup>Hf = 0.2%, <sup>176</sup>Hf/<sup>177</sup>Hf = 0.005%. Minerals are grt, garnet; cpx, clinopyroxene; wr, whole rock.

<sup>b</sup>Sample weights are listed in mg.

<sup>c</sup>The  $\epsilon_{\text{Hf}(0)}$  was calculated using present-day ratios of <sup>176</sup>Hf/<sup>177</sup>Hf = 0.282772 and <sup>176</sup>Lu/<sup>177</sup>Hf = 0.0334 for CHUR [Blichert-Toft and Albarède, 1997].

<sup>d</sup>Purer garnet separate.

<sup>e</sup>Less pure garnet separate.

California, Santa Barbara. All quoted age uncertainties are  $2\sigma$  and exclude uncertainty in decay constants.

### 3.1. Sample Descriptions

[10] Sample P5701A from Sandvik, northern Gurskøy (Figure 1) is an eclogite that contains little evidence for retrogression. The sample consists of agglomerations of hypidioblastic to xenoblastic garnet 0.1–1 mm in diameter (~55 vol %). Clinopyroxene (40 vol %) up to 0.5 cm wide and 2 cm long defines a millimeter-scale foliation and has minor amphibole replacement. Biotite (<5 vol %) is ~1 mm long and aligned parallel to the foliation. Minor opaque minerals occur along grain boundaries. The nearest eclogite facies pressure and temperature estimates are 795°C and 3.2 GPa from Hareidlandet, ~10 km to the NE [Carswell *et al.*, 2003b].

[11] Sample E1612Q5 is an eclogite from Hellesylt (Figure 1) [Walsh *et al.*, 2007]. It contains millimeter-scale layers of hypidioblastic garnet 0.1–1 mm in diameter and ~2-mm-long omphacite. Rutile (<2 vol %) occurs as isolated ~0.2-mm grains. Pressures of 2.8 GPa and temperatures >700°C have been estimated from samples within a few kilometers [Walsh and Hacker, 2004].

[12] Sample NOR205 is an unretrogressed eclogite from the Blåhø Nappe on Gossa (Figure 1), described by Hollocher *et al.* [2007]. Garnet (~30 vol %) is 2–4 mm and poikiloblastic, containing clinopyroxene, calcite, and amphibole. Clinopyroxene (~20 vol %) is 1–5 mm and exhibits strong undulatory extinction and subgrains. Orthopyroxene (~20 vol %) is ~2–5 mm and has few subgrains. Amphibole (~20 vol %) is 0.2 mm and hypidioblastic. Minor phases include rutile and opaque minerals (<2 vol

%). Pressure-temperature estimates from nearby UHP rocks are ~800°C and ~3.3 GPa [Carswell *et al.*, 2006].

[13] Sample K5622A1 is an unfoliated and unretrogressed eclogite from Vollstein (Figure 1) composed of equal portions of clinopyroxene (~0.2–1 mm) and garnet. Garnet is idioblastic, ~0.5 mm, and has darker pink cores with inclusions of amphibole and epidote. Amphibole (~10 vol %) is xenoblastic and ~0.2–1 mm in diameter. Minor phases include phengite, rutile, and epidote. Thermobarometry using garnet-clinopyroxene-phengite equilibria [Krogh Ravn and Terry, 2004] produced temperatures and pressures of 615°C and 2.4 GPa.

### 3.2. Lu-Hf and Sm-Nd Ages

[14] Our new Lu-Hf and Sm-Nd data are presented in Table 1a and 1b and Figure 4. All the samples yielded relatively low MSWD values ( $\leq 2.5$ ) for all fractions analyzed. One sample generated a high-precision three-point Lu-Hf isochron, and three samples generated relatively high-precision four-point Sm-Nd isochrons. Two samples for which we obtained both Lu-Hf and Sm-Nd ages (E1612Q5 and P5701A) yielded relatively low parent/daughter ratios for the garnet fractions, and thus low-precision ages. This is likely the result of a relatively low Lu/Hf ratio in the whole rock and/or a larger modal proportion of garnet, rather than the presence of high-Hf inclusions in the garnet fractions. Given the high distribution coefficient for Lu in garnet, the bulk of the Lu is likely sequestered in garnet cores, whereas the Hf concentration should be relatively constant from core to rim [Lapen *et al.*, 2003; Skora *et al.*, 2006]. Thus, the larger the modal proportion of garnet, or the smaller the Lu/Hf ratio in the whole rock, the smaller the Lu/Hf ratio in the bulk garnet

**Table 1b.** Samarium-Neodymium Isotope Data for Eclogites From the Western Gneiss Region, Norway<sup>a</sup>

Sample <sup>b</sup>	Concentration (ppm)		<sup>147</sup> Sm/ <sup>144</sup> Nd	<sup>143</sup> Nd/ <sup>144</sup> Nd	Age (Ma)	$\epsilon_{Nd(0)}$ <sup>c</sup>
	Sm	Nd				
P5701A (Sandvik)						
grt (86.8)	0.93	0.47	1.205	0.515043 ± 12		
cpx (53.3)	3.27	10.91	0.182	0.512365 ± 13		
wr (44.7)	1.75	5.92	0.179	0.512383 ± 9	397.1 ± 4.8	-4.3
E1612Q5 (Geiranger)						
grt <sup>d</sup> (70.7)	1.31	0.95	0.837	0.514085 ± 9		
grt <sup>e</sup> (93.4)	1.31	1.01	0.792	0.513986 ± 8		
cpx (37.0)	1.91	5.68	0.204	0.512467 ± 10		
wr (183.4)	1.98	5.55	0.216	0.512452 ± 12	398.3 ± 8.1	-4.1
NOR205 (Gossa)						
grt <sup>3</sup> (37.3)	1.26	0.59	1.301	0.515909 ± 11		
grt <sup>4</sup> (33.6)	1.17	0.57	1.250	0.515789 ± 14		
cpx (35.9)	0.78	2.47	0.191	0.512906 ± 14		
wr (223.7)	1.59	3.80	0.254	0.513083 ± 7	413.9 ± 3.7	5.6

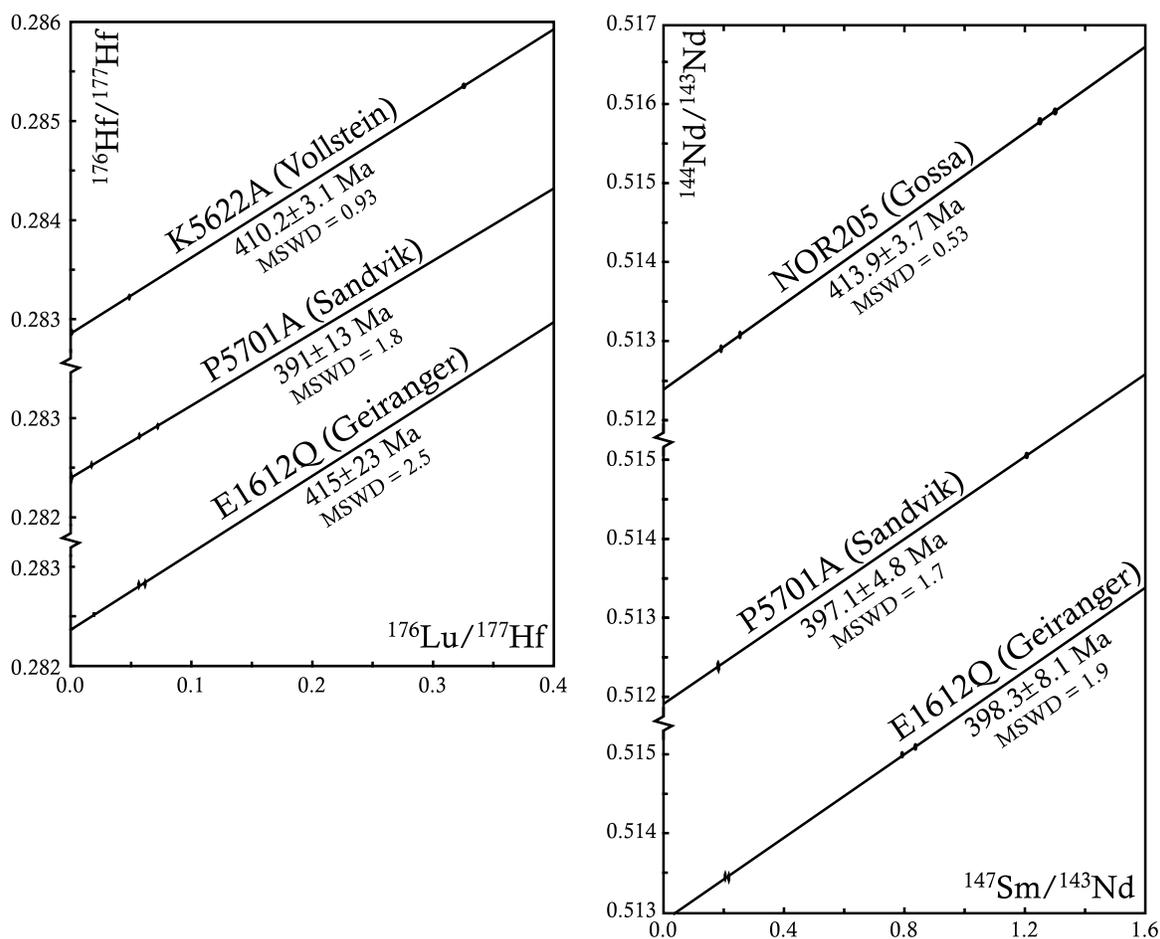
<sup>a</sup>Analysis accomplished by isotope dilution multicollector plasma mass spectrometry at University of Wisconsin, Madison. Decay constant used for <sup>147</sup>Sm is  $6.54 \times 10^{-12} \text{ a}^{-1}$ . Errors in Nd isotope ratios are expressed as 2SE and refer to least significant digits. Isochron ages were calculated using Isoplot v. 3.0. Errors calculated for ages are based solely on external reproducibility of spiked standards and whole rock samples (errors for individual analyses are negligible): <sup>147</sup>Sm/<sup>144</sup>Nd = 0.5%, <sup>143</sup>Nd/<sup>144</sup>Nd = 0.005%.

<sup>b</sup>Sample weights are listed in mg.

<sup>c</sup>The  $\epsilon_{Nd(0)}$  was calculated using present-day ratios of <sup>146</sup>Nd/<sup>144</sup>Nd = 0.7219 and <sup>143</sup>Nd/<sup>144</sup>Nd = 0.512638 for CHUR.

<sup>d</sup>Purer garnet separate.

<sup>e</sup>Less pure garnet separate.

**Figure 4.** New Sm-Nd and Lu-Hf ages from the WGR, Norway. Sample locations in Figure 1.

separate. Our previous Lu-Hf data set for WGR eclogites produced this same phenomenon [Kylander-Clark *et al.*, 2007]. Nevertheless, any significant volume of high-Hf inclusions could not have significantly affected the age yielded by both samples. If high-Hf inclusions, such as zircon, were significantly older than the enclosing garnet, the inclusion-rich garnet fraction would have lower  $^{176}\text{Hf}/^{177}\text{Hf}$  ratios relative to the inclusion-poor garnet fraction [Prince *et al.*, 2000]; this is not the case in either of our samples. Furthermore, variable amounts of old zircon in each of the handpicked separates would yield data that are not isochronous.

[15] The garnet ages span more than 20 Ma, from 415 Ma to 391 Ma. Because of the large uncertainty for the youngest age, a more realistic span of  $\sim 17 \pm 6$  Ma is given by the difference between two high-precision ages of  $413.9 \pm 3.7$  Ma and  $397.1 \pm 4.8$  Ma, where the uncertainty in the span in ages is estimated using the square root of the sum of the squares of the errors of the two ages. The Sm-Nd age from Gossa (NOR205),  $413.9 \pm 3.7$  Ma, is the northernmost four-point isochron obtained from the WGR; it is equivalent to Sm-Nd garnet, U-Pb zircon and U-Pb monazite ages from nearby localities (Figure 1) [Griffin and Brueckner, 1980; Terry *et al.*, 2000a; Krogh *et al.*, 2004]. The Sm-Nd age of  $398.3 \pm 8.1$  Ma obtained from Hellesylt (E1612Q5) is equivalent to the  $396 \pm 10$  Ma U-Pb zircon SIMS age obtained by Walsh *et al.* [2007] for the same sample. The Lu-Hf age from E1612Q5 overlaps its Sm-Nd age, but errors are high because there is little spread in Lu/Hf ratios. Sample P5701A from north Gurskøy yielded a high-precision Sm-Nd age of  $397.1 \pm 4.8$  Ma, similar to the Sm-Nd age of  $398.3 \pm 5.5$  Ma from west Gurskøy [Kylander-Clark *et al.*, 2007]. The Lu-Hf age of  $391 \pm 13$  Ma for this sample lies within error of its Sm-Nd age, although the errors for the Lu-Hf age are relatively high because of the limited spread in Lu/Hf ratios. The Lu-Hf age of  $410.2 \pm 3.1$  from sample K5622A overlaps Lu-Hf ages from the central WGR [Kylander-Clark *et al.*, 2007] and is the first eclogite age obtained from the southern WGR (Figure 1).

### 3.3. Discussion of the Geochronology Data Set

[16] Our new ages, combined with the existing geochronologic data set for the WGR, strongly suggest that the (U)HP metamorphism of the WGR occurred over a 20 Ma period: U-Pb zircon and monazite ages for eclogites at (U)HP conditions range from 415 to 400 Ma, high-precision Lu-Hf ages range from 420 to 410 Ma, and high-precision Sm-Nd ages range from 412 to 397 Ma (Figure 1). This conclusion is reinforced by the fact that this data set comes from three different isotopic systems from measurements made in different laboratories.

[17] These eclogite ages come from a large portion of the WGR, spanning  $\sim 220$  km north to south within the UHP hinterland and 60 km from the hinterland in the west toward the foreland in the east. Does such a long span of ages mean that this entire area was subjected to (U)HP metamorphism for 20 Ma as a large block, or is it possible that this large region comprises different slices that had distinct *P-T* histories and hence underwent (U)HP metamorphism at

different times? There are two observations that argue in favor of the former and against the latter.

[18] 1. There are no UHP domains that have yielded a narrow range of eclogite ages. For example, the Nordfjord UHP domain yields a Lu-Hf age at Verpeneset  $\sim 20$  Ma older than the U-Pb zircon age of Flatraket, only 8 km distant. The Nordøyane UHP domain has produced U-Pb and Sm-Nd ages of 415–410 Ma, but also a Sm-Nd age of  $388 \pm 10$  Ma (Figure 1). It is possible, albeit unlikely, that the young age does not represent eclogite facies conditions [Kylander-Clark *et al.*, 2007] and thus could indicate that the northern WGR underwent a different prograde history than the WGR south of the Nordøyane UHP domain. An amphibolite facies shear zone that cuts through the southern portion of this region [Terry and Robinson, 2003, 2004] would support this hypothesis, but it separates the  $\sim 412$  Ma Løvsøya eclogite from similar age eclogites to the north [Terry *et al.*, 2000a; Krogh *et al.*, 2004]. No other significant structures have been recognized.

[19] 2. Two eclogite localities, Vigra and NW Gurskøy (Figure 1), have Lu-Hf ages that are 14 Ma older than their respective Sm-Nd ages, indicating that garnet growth spanned at least that long [Kylander-Clark *et al.*, 2007]. Thus, the  $\sim 20$  Ma spread in ages throughout the WGR is best explained by long-term residence of a  $\sim 220$  km  $\times$   $\sim 60$  km piece of subducted continental crust at high-pressure conditions.

[20] Although the (U)HP eclogite facies metamorphism was preceded and followed by metamorphic episodes in which garnet was stable, it is unlikely that the garnet ages reflect mixtures of garnet of a variety of ages. First, the dated samples do not contain significant retrograde mineral assemblages. Second, although a Barrovian amphibolite facies metamorphism at  $\sim 700^\circ\text{C}$  and 1.0–1.5 GPa occurred between 445 and 432 Ma over much of the allochthons in the foreland [Hacker and Gans, 2005], this event has not been recognized in the crystalline WGC. Third, although Precambrian ( $\sim 950$ – $900$  Ma) granulite facies metamorphism is recognized in a small portion of the crystalline rocks of the WGR [Cohen *et al.*, 1988; Røhr *et al.*, 2004; Root *et al.*, 2005; Glodny *et al.*, 2008], and in some places is overprinted by eclogite facies metamorphism [Wain *et al.*, 2001], these two events produced distinctly different major element zoning in garnets. The samples studied here contain no evidence of granulite facies garnet compositions. In addition, because the granulite facies metamorphism predated the eclogite facies metamorphism by 500 Ma, a sample that contains a mixed population of garnets should not produce a multipoint, low-MSWD isochron typical of our data. Furthermore, Caledonian garnets that contain an inherited component of  $\sim 950$  Ma garnet could be no more than 5% inherited to yield an age between 425 and 400 Ma. All of the 16 sample locations listed in Figure 1 that report either a Lu-Hf age, Sm-Nd age, or both, yield ages between 425 and 400 Ma. Because it is no more likely for a garnet to contain 5% inheritance than, for example, 50% inheritance, the span in ages likely represents the minimum duration of HP conditions during the Caledonian event.

[21] Some eclogites in the WGR had a prograde history that passed through amphibolite or blueschist facies conditions, during which time garnet growth likely occurred [Krogh, 1980, 1982] at pressures of  $\sim 1$  GPa and higher. Because the pressures of eclogite facies metamorphism in the WGR ranged up to 3.4 GPa [Terry *et al.*, 2000a; Krogh Ravn and Terry, 2004; Carswell *et al.*, 2006], garnet growth probably occurred at depths of  $\sim 33$  km to 112 km (assuming  $3 \text{ g cm}^{-3}$ ). The inferred  $\sim 20$  Ma interval of garnet growth, therefore, is interpreted to reflect prograde growth during burial at an average rate of  $<4 \text{ mm a}^{-1}$ .

#### 4. Thermal Model of Long-Duration (U)HPM Tectonism

[22] Although it is generally assumed that UHP terranes spent only a few million years at UHP conditions, at least four regions appear to have spent 15–20 Ma at such conditions: the Norwegian UHP terrane (this paper), the Dabie-Sulu terrane of eastern China [Hacker *et al.*, 2006], the North Qaidam terrane of western China [Mattinson *et al.*, 2006], and the northeast Greenland eclogite province [McClelland *et al.*, 2006]. Long-term metamorphism at UHP conditions is a puzzle for one simple reason: the characteristic thermal diffusion distance for 15–20 Ma is 20–25 km (assuming a thermal diffusivity of  $10^{-6} \text{ m s}^{-2}$  [Turcotte and Schubert, 2002]), suggesting that if UHP terranes are only a few kilometers thick, they cannot survive such a lengthy immersion in the hot mantle without melting [Root *et al.*, 2005]. In this section we use a simple thermal model to explore the parameter space (i.e., subduction rate and initial thermal profile) that might permit terranes (1) to survive 15–20 Ma at HP conditions without exceeding the maximum temperatures observed in the WGR and (2) to undergo near-isothermal return to midcrustal levels.

##### 4.1. Model Parameters

[23] The thermal model is a  $395 \times 195$  km (5-km node spacing) 2-D transient heat flow model that uses the alternating direction, implicit, finite difference method described by Hacker [1990]. The upper plate is continental and the lower plate is continental or oceanic (see below). The continental crust is 40 km thick and has a radiogenic heat production rate of 1.68, 1.04, and  $0.21 \mu\text{W m}^{-3}$  in the upper (top 12 km), middle (12–23 km), and lower (23–40 km) crust, respectively (average for continental crust [Rudnick and Gao, 2003]). The upper boundary of the model is held at  $0^\circ\text{C}$ , the basal boundary condition is a constant heat flow of  $25 \mu\text{W m}^{-2}$ , and the lateral boundary conditions are a downward increasing constant temperature gradient. The angle of subduction of the lower plate is  $45^\circ$ . Convection in the mantle wedge is prescribed in a simplified form; corner flow occurs at a constant flow rate everywhere the temperature exceeds  $1000^\circ\text{C}$  (Figure 5a). The time step was 0.01 Ma for subduction and 0.005 Ma for exhumation.

[24] The two primary controls on the thermal structure of a subducting plate are subduction velocity and initial thermal structure [e.g., van Keken *et al.*, 2002]. We model

two end-member scenarios: (1) subduction of a continental lower plate into a thermal steady state, Andean-style subduction zone and (2) intracontinental subduction. In both scenarios we vary the continental subduction rate and initial thermal gradient to explore the effects on the thermal evolution of the system. In scenario 1 we use an upper plate thermal gradient of  $15^\circ\text{C km}^{-1}$ , similar to the thermal structure of Andean-style subduction zones [van Keken *et al.*, 2002]. For simplicity, the subducting oceanic plate is given the thermal structure of the subducting continental plate. The oceanic subduction velocity is  $47 \text{ mm a}^{-1}$ , equivalent to the average subduction rate of Andean-style subduction zones [Jarrard, 2003; Lallemand *et al.*, 2005] and the model is run for 30 Ma to reach steady state thermal conditions. The model is then run with the prescribed continental subduction velocity until the continental material reaches a depth of  $\sim 105$  km ( $\sim 3.2$  GPa) or 150 km ( $\sim 5$  GPa).

[25] The initial thermal gradients for the continental material were varied from  $10^\circ\text{C km}^{-1}$  to  $25^\circ\text{C km}^{-1}$ . The vertical component of the subduction velocity was varied between 1 and  $21 \text{ mm a}^{-1}$ , a range that encompasses all modern convergence zones with continental lower plates. Maintenance of all points in the subducted continent within the garnet-stable depth range requires subduction rates of  $<4 \text{ mm a}^{-1}$  for metamorphic durations of 20 Ma.

[26] After subduction, the motion of the upper 40 km of the lower plate was reversed to simulate 85 km of exhumation from a maximum depth of 105 km to a midcrustal depth of 20 km (Figure 5b). The vertical exhumation rate was  $33 \text{ mm a}^{-1}$ , corresponding to 85 km of exhumation in 5 Ma, as suggested by geochronology (see section 3.3).

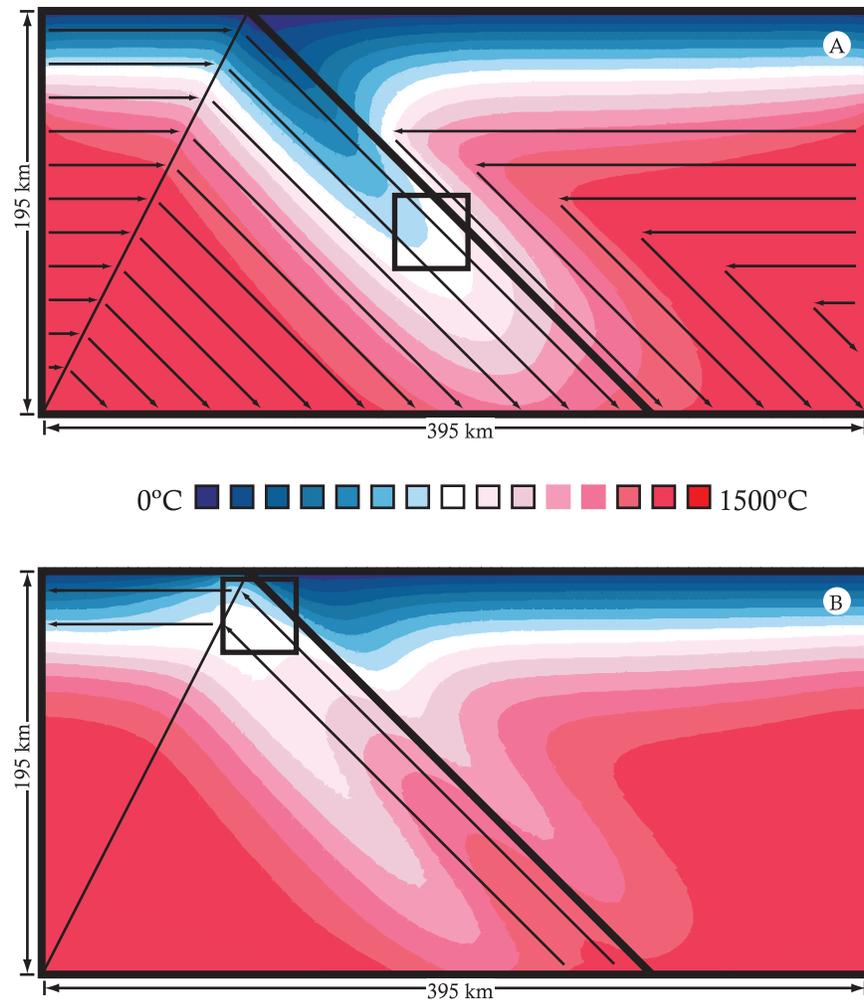
##### 4.2. Model Results

[27] Examples of the calculated temperature field are shown in Figure 6a. As expected, temperatures in the core of the subducted slab are relatively low if the initial slab thermal gradient is low (Figure 6a, left) or the subduction rate is high (Figure 6a, right). Only for vertical subduction rates of  $\sim 1$ – $11 \text{ mm a}^{-1}$  ( $\sim 2$ – $16 \text{ mm a}^{-1}$  subduction at  $45^\circ$ ) and slab thermal gradients of  $10$ – $25^\circ\text{C km}^{-1}$  do the temperatures at 105 km depth match the  $700$ – $800^\circ\text{C}$  temperatures observed in Norway. During exhumation, temperatures at the top of the lower plate always decrease, generally by more than  $200^\circ\text{C}$ . Temperatures at depth in the slab increase for fast subduction ( $\geq 16 \text{ mm a}^{-1}$ ), and decrease for slow subduction ( $\leq 11 \text{ mm a}^{-1}$ ).

[28] Figure 6b shows examples of the temperature field in the crust after exhumation. Only for vertical subduction rates of  $<10 \text{ mm a}^{-1}$  ( $<15 \text{ mm a}^{-1}$  subduction at  $45^\circ$ ) do the models produce the near-isothermal exhumation to midcrustal depths observed in Norway. Figure 6c shows  $T_d$ , the sum of the temperature deviations at each node from  $750^\circ\text{C}$  at peak pressure and after exhumation, calculated as

$$T_d = |750^\circ\text{C} - T_{\text{at\_peak\_pressure}}| + |750^\circ\text{C} - T_{\text{at\_15\_km}}|$$

The white domains of Figure 6c indicate conditions that produce near-isothermal paths; cooler colors show UHP



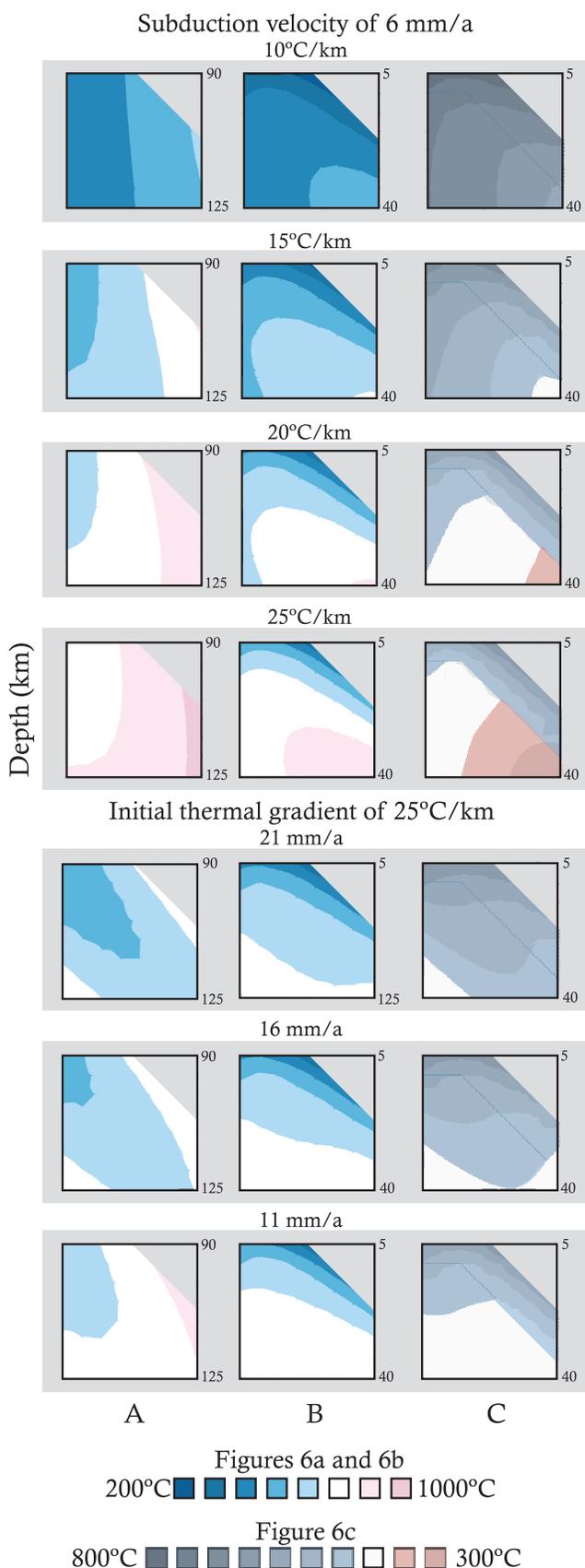
**Figure 5.** Flow lines for thermal model showing crustal sections represented in Figure 6 and used to make calculations shown in Figure 7 (black squares). (a) Subduction. (b) Exhumation. Convection in the mantle wedge is dependent on initial thermal gradient and is defined in this model at  $\sim 1000^{\circ}\text{C}$ . The temperatures shown in this example result from intracontinental subduction with an initial thermal gradient of  $25^{\circ}\text{C km}^{-1}$  and a subduction velocity of  $11 \text{ mm a}^{-1}$ .

metamorphism that is colder than  $750^{\circ}\text{C}$ , or cooling during exhumation, whereas warmer colors show UHP metamorphism hotter than  $750^{\circ}\text{C}$  (slow subduction), or warming during exhumation (fast subduction). The cooler colors at the top and upper left of each diagram reflect excessive cooling, as the crust has reached shallow levels and been cooled from above.

[29] Figure 7 summarizes the average temperature and temperature deviation,  $T_d$ , for crust that reaches 105 to 90 km depth. Only the middle 28 km of the crust is shown; the outer portion of the crust becomes too hot during subduction and/or too cold during exhumation to be representative of the exposed terrane in Norway, as discussed in section 4.3.1. For 105 km of subduction, only intracontinental subduction is shown because the Andean subduction scenario only yields temperatures higher than  $700^{\circ}\text{C}$  if

subduction rates are unrealistically slow ( $\sim 1 \text{ mm a}^{-1}$  or less). From these considerations, we infer that vertical subduction rates must be slow ( $\leq 4 \text{ mm a}^{-1}$ ; equivalent to  $\leq 6 \text{ mm a}^{-1}$  subduction velocity) to maintain the subducted continent in the garnet stability field (30 to 60–105 km depth) for 20 Ma, as well as produce  $750^{\circ}\text{C}$  UHP eclogites during subduction and near-isothermal exhumation. The main difference between the Andean and the intracontinental subduction scenarios is that continental subduction in the Andean setting yields cooler temperatures than intracontinental subduction, for any given set of initial conditions.

[30] The largest portions of crust undergo near-isothermal decompression at  $\sim 750^{\circ}\text{C}$  (lowest  $T_d$  values) for subduction to either 105 or 150 km depth if the slab is initially hot (pale colors in Figure 7c). The Andean subduction scenario achieves minimum  $T_d$  values for 150 km of subduction



and exhumation at subduction rates of  $<2 \text{ mm a}^{-1}$ ; this rate and depth require  $>75 \text{ Ma}$  of subduction, which is incompatible with the Norway geochronology data set. The intracontinental subduction scenario yields the lowest  $T_d$  values at subduction rates of  $\sim 6$  and  $11 \text{ mm a}^{-1}$  for 150 and 105 km of subduction, respectively. For vertical subduction rates of  $2\text{--}4 \text{ mm a}^{-1}$ , the same scenario yields minimum  $T_d$  values for moderate initial thermal gradients ( $15\text{--}20^\circ\text{C km}^{-1}$ ) for subduction to 105 km and higher initial thermal gradients ( $20\text{--}25^\circ\text{C km}^{-1}$ ) for subduction to 150 km. Intracontinental subduction also leads to larger variation in average temperature and  $T_d$  values, reflecting the greater variation in the initial upper plate temperature profiles. These scenarios are compatible with the Norway geochronology database.

### 4.3. Discussion of the Modeling

#### 4.3.1. Implications

[31] Because rocks over a large region of the WGR followed similar  $T$ - $t$  paths (Figure 2) and recorded peak pressures mostly  $\leq 3.2 \text{ GPa}$ , model runs that yield the lowest  $T_d$  values produce  $P$ - $T$ - $t$  paths that are closest to those observed in Norway. Coupled with the constraints that (1) vertical subduction rates of  $2\text{--}4 \text{ mm a}^{-1}$  are required by the  $\sim 20\text{-Ma}$  duration of WGR UHP metamorphism and (2) decompression was rapid and nearly isothermal, our modeling results favor an intracontinental setting in which the slab had a moderate to hot geothermal gradient ( $15\text{--}25^\circ\text{C km}^{-1}$ ) and was subducted at a vertical rate of  $\sim 2\text{--}4 \text{ mm a}^{-1}$ . The models do not preclude cooler thermal gradients or Andean-style subduction, but on the basis of the model results, these possibilities seem less likely.

[32] The simulations also make some useful predictions.

[33] 1. Rocks that were farther up dip of the UHP domain during subduction, should have reached lower temperatures and decompressed isothermally. In Norway, eclogites that record lower temperatures and pressures are east of the UHP domains [Cuthbert *et al.*, 2000; Young *et al.*, 2007a] and are overprinted by amphibolite facies assemblages that formed at temperatures similar to the eclogite facies temperatures [Walsh and Hacker, 2004].

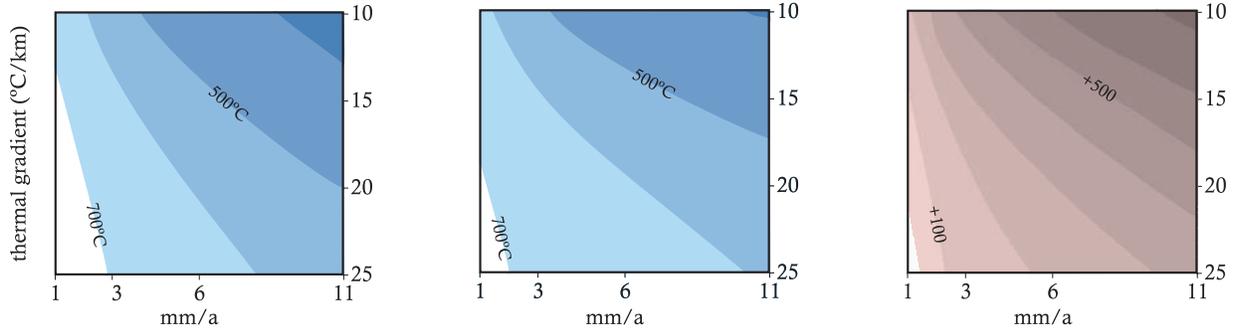
[34] 2. Rocks that were farther down dip of the UHP domain should have reached higher temperatures. Such rocks are yet unrecognized in the WGR and may not be exposed because of rifting of the Atlantic Ocean.

**Figure 6.** (a) Temperatures of the crust, in a  $35 \times 35 \text{ km}$  section (black square from Figure 5a) from 90 to 105 km following 150 km of intracontinental subduction, using varied initial thermal gradients and subduction rates. (b) Exhumation of each crustal section in Figure 6a, after 85 km of exhumation. (c) Deviation from isothermal decompression at  $750^\circ\text{C}$  following exhumation. Warm colors show excessive temperatures upon subduction or exhumation; cooler colors show cold initial temperatures or excessive cooling. The upper plate (top right) is shaded out in each diagram for simplicity.



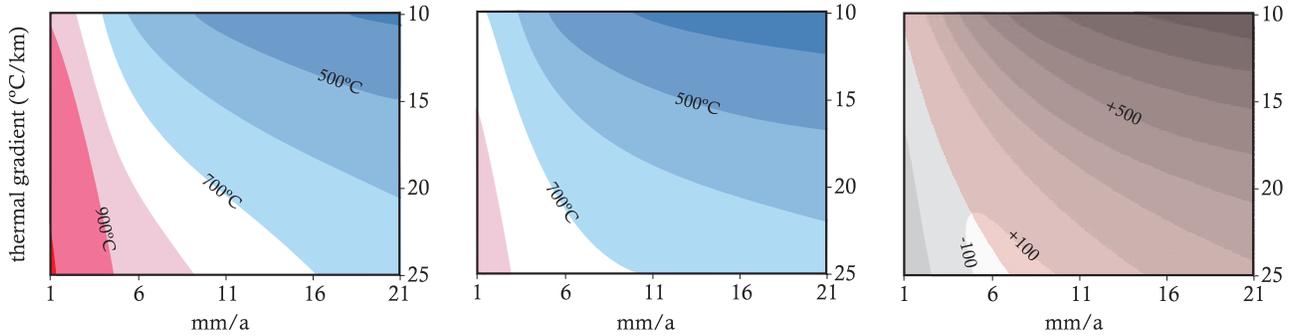
continental subduction following Andean style subduction

150 km subduction from 105 km

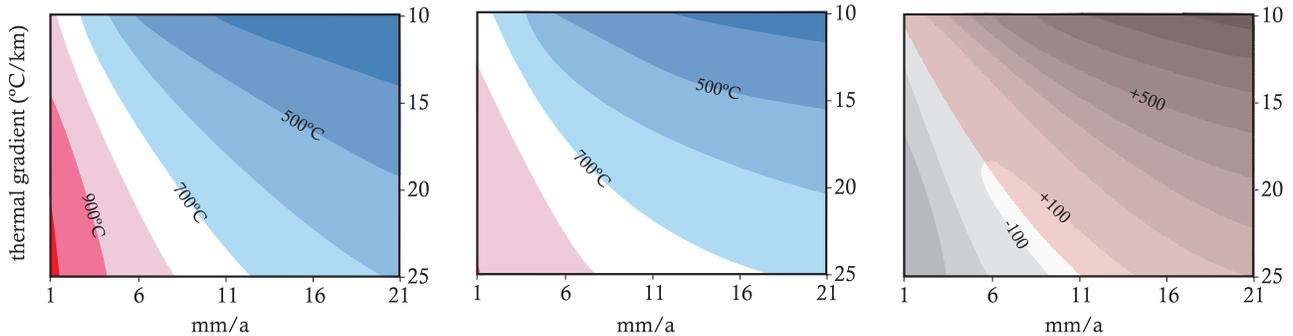


intracontinental subduction

150 km subduction from 105 km



105 km subduction from 105 km



A. subduction

B. exhumation

C. deviation

**Figure 7.** (a) Average temperatures for the middle ~28 km of the crust, from 90 to 105 km following 150 or 105 km of subduction, with variable subduction velocities and initial thermal gradients of the lower plate. (b) The portion of crust in Figure 7a, after 85 km of exhumation. (c) Deviation from isothermal decompression at 750°C of the crustal section represented in Figures 7a and 7b.

[35] 3. The downdip thermal gradient in satisfactory scenarios ranges from ~3 to 5°C km<sup>-1</sup>, similar to that observed in Norway [e.g., Young *et al.*, 2007a].

**4.3.2. Limitations**

[36] There are several simplifications in our model that may produce uncertainties relative to the thermal-mechan-

ical conditions that are expected during UHP metamorphism. These uncertainties are discussed in detail below; however, our model yields similar *P-T* paths of more complicated thermal-mechanical models [Warren *et al.*, 2008], given similar initial conditions.

[37] A constant subduction velocity was assumed, and a single upper plate thermal gradient, similar to that produced by recent thermal models [e.g., *van Keken et al.*, 2002], was used to establish the thermal profile of the Andean-style margin prior to continental subduction. The thermal profile of the upper plate is essentially independent of subduction velocity and the age of the lower plate. The thermal profile of the upper plate does, however, depend on parameters, such as crustal thickness, that are not well known for Norway. Therefore, the initial thermal profile of the upper plate could have been significantly different from our model. For example, an increase of the upper plate thermal gradient by  $10^{\circ}\text{C km}^{-1}$  yields an increase of  $\sim 100^{\circ}\text{C}$  in the lower plate at 100 km depth; such an increase in temperature allows a large portion of the crust to reach  $750^{\circ}\text{C}$  and to decompress isothermally, following reasonable vertical rates of subduction (i.e.,  $2\text{--}4\text{ mm a}^{-1}$ ).

[38] Our fixed subduction angle of  $45^{\circ}$  is steeper than some active continental convergence zones (e.g., India-Asia), but shallower than others (e.g., Hindu Kush). Nevertheless, our models compare favorably to those that assumed shallower dips, such as *Roselle and Engi* [2002], who used a subduction angle of  $30^{\circ}$ . For the same vertical subduction rate and initial thermal conditions, our model produces similar temperatures at UHP depths.

[39] Frictional heating is poorly understood in subduction zones and ignored in our models. Incorporation of friction in the model would produce higher temperatures in the lower plate, with a greater effect at faster plate velocities. The heating due to friction is highest at the plate boundary and typically no greater than  $100^{\circ}\text{C}$  [e.g., *Roselle and Engi*, 2002]. Because the plate boundary cools by  $>200^{\circ}\text{C}$  in our models, we anticipate that crustal cooling in excess of  $100^{\circ}\text{C}$  would still occur during exhumation if frictional heating were included. The uncertainties in frictional heating may be mitigated if we only consider the middle 28 km of the subducted crust in our discussion because it is anticipated that the middle section of crust is the most likely portion to experience isothermal decompression.

[40] We assume a simple convection model for the mantle wedge. More complex models for the upper plate have been used in other studies [e.g., *van Keken et al.*, 2002], but uncertainties in the subduction geometry during formation of the Norwegian Caledonides prevent us from exploring these. We note that our models do produce results that are similar to those obtained using more complicated models for the upper plate [e.g., *Roselle and Engi*, 2002].

[41] The subducted plate is assumed a rigid block during subduction and exhumation. The plate could thicken or thin during subduction and/or exhumation, and this deformation would affect the temperature within the slab. For example, given the same vertical subduction rate, a slab that thickened during subduction would have a colder core than a slab that did not thicken. Likewise, thinning during exhumation would allow the slab to cool more rapidly.

[42] Finally, we assume constant subduction and exhumation rates. Long-term HP conditions can be produced by temporal variation in subduction rate, such as faster subduction to near-UHP depths, followed by slower subduction

at UHP depths. Likewise, a 5 Ma decompression path can occur via temporal variations in exhumation rates, such as rapid exhumation through the mantle, followed by slower exhumation through the lower crust. The more time the material in question spends at greater depths, the hotter it will be than those represented in our model.

## 5. Discussion of Tectonics

[43] The modeling discussed in section 4.3 suggests that HP conditions can be maintained for  $\sim 20$  Ma, followed by near-isothermal decompression at  $750^{\circ}\text{C}$ , provided that (1) relatively slow subduction occurred in an intracontinental setting and (2) exhumation involved a crustal segment that was at least tens of kilometers in the minimum dimension. Subduction rates for UHP terranes as slow as a few millimeters per year were formerly thought to be unlikely because the crustal segments that went to UHP conditions were envisioned to be thin to facilitate the rapid cooling observed in UHP terranes worldwide [*Ernst and Peacock*, 1996]; during slow subduction, thin crustal sections would heat too much during subduction. In contrast, *Root et al.* [2005] explained the  $P$ - $T$  history of the WGR, which remained at  $\sim 750^{\circ}\text{C}$  from the time of UHP metamorphism to the time it reached midcrustal depths of  $\sim 15\text{--}20$  km [*Terry et al.*, 2000a; *Root et al.*, 2005], through subduction of a UHP body  $>30$  km thick. Our more detailed modeling supports this conclusion and emphasizes the observation that a more complex simulation of exhumation leads to unacceptably large cooling in at least the outer 7 km of a 40-km-thick slab (more for slower exhumation rates). We conclude that the slab must have been more than 14 km thick to allow isothermal cooling during exhumation. *Warren et al.*'s [2008] more complicated model also shows an exhumed UHP wedge that is  $>20$  km thick. There are few geologic observations that constrain the thickness of the WGR. Seismic studies indicate that western edge of the WGR consists of a  $>40$ -km-thick section of exhumed Baltic crust, which supports our findings; however, this section may consist of imbricate slices [*Hurich*, 1996].

[44] Following near-isothermal exhumation to midcrustal depths of 20 km, the WGR cooled rapidly. Specifically, the cooling rate across the WGR varied from  $15$  to  $50^{\circ}\text{C Ma}^{-1}$  ( $\sim 750^{\circ}\text{C}\text{--}400^{\circ}\text{C}$  in  $5\text{--}15$  Ma), with faster cooling in the eastern WGR. Because the characteristic thermal diffusion distance for  $5\text{--}15$  Ma is  $12\text{--}21$  km, this cooling could have been accomplished by conductive cooling from above.

[45] We conclude that UHP terranes that are large and contain evidence for a protracted period of peak or near-peak metamorphism represent one end-member of a spectrum of size–time–subduction rate parameters. We suggest that other UHP terranes that show similar traits over such large areas also may have been thick bodies from the onset of subduction through exhumation to lower to midcrustal levels. An obvious candidate is the Dabie-Sulu terrane [*Liou et al.*, 1997; *Zhang et al.*, 2003; *Hacker et al.*, 2006; *Liu et al.*, 2006a, 2006b]. At the other end of the spectrum, smaller, thinner UHP units, such as the Kaghan Valley terrane [*Kaneko et al.*, 2003], that reached high temper-

atures and retained them through the bulk of the exhumation, probably followed a rapid subduction, rapid exhumation path, where the terranes spent minimal time at peak or near-peak metamorphic conditions.

## 6. Conclusions

[46] New high-precision Lu-Hf and Sm-Nd ages for (U)HP eclogites from the Western Gneiss Region of Norway from  $413.9 \pm 3.7$  to  $397.1 \pm 4.8$  Ma add to existing data that document a 20-Ma record of high-pressure metamorphism for an area spanning  $60 \times 220$  km. These data require slow average subduction ( $\sim 2\text{--}4$  mm a<sup>-1</sup>) and rapid,

near-isothermal exhumation ( $\sim 10\text{--}30$  mm a<sup>-1</sup>), followed by rapid cooling ( $15\text{--}50^\circ\text{C Ma}^{-1}$ ). Thermal models indicate that this *P-T-t* path is likely to have occurred at  $>7$  km depth within a downgoing slab  $>30$  km thick during subduction that initiated within a relatively warm continent. Other UHP terranes that show similar geochronologic and *P-T* characteristics, such as the Dabie-Sulu UHP terrane, may have developed in a similar manner.

[47] **Acknowledgments.** Samples of the Gossa eclogite were generously provided by K. Hollocher, P. Robinson, and E. Walsh, who collected the sample in 2004. The eclogite was discovered in 2000 by P. Robinson, and S. Cuthbert. R. Parrish offered an insightful review. Funding provided by NSF grant EAR-0510453 to B. R. Hacker.

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